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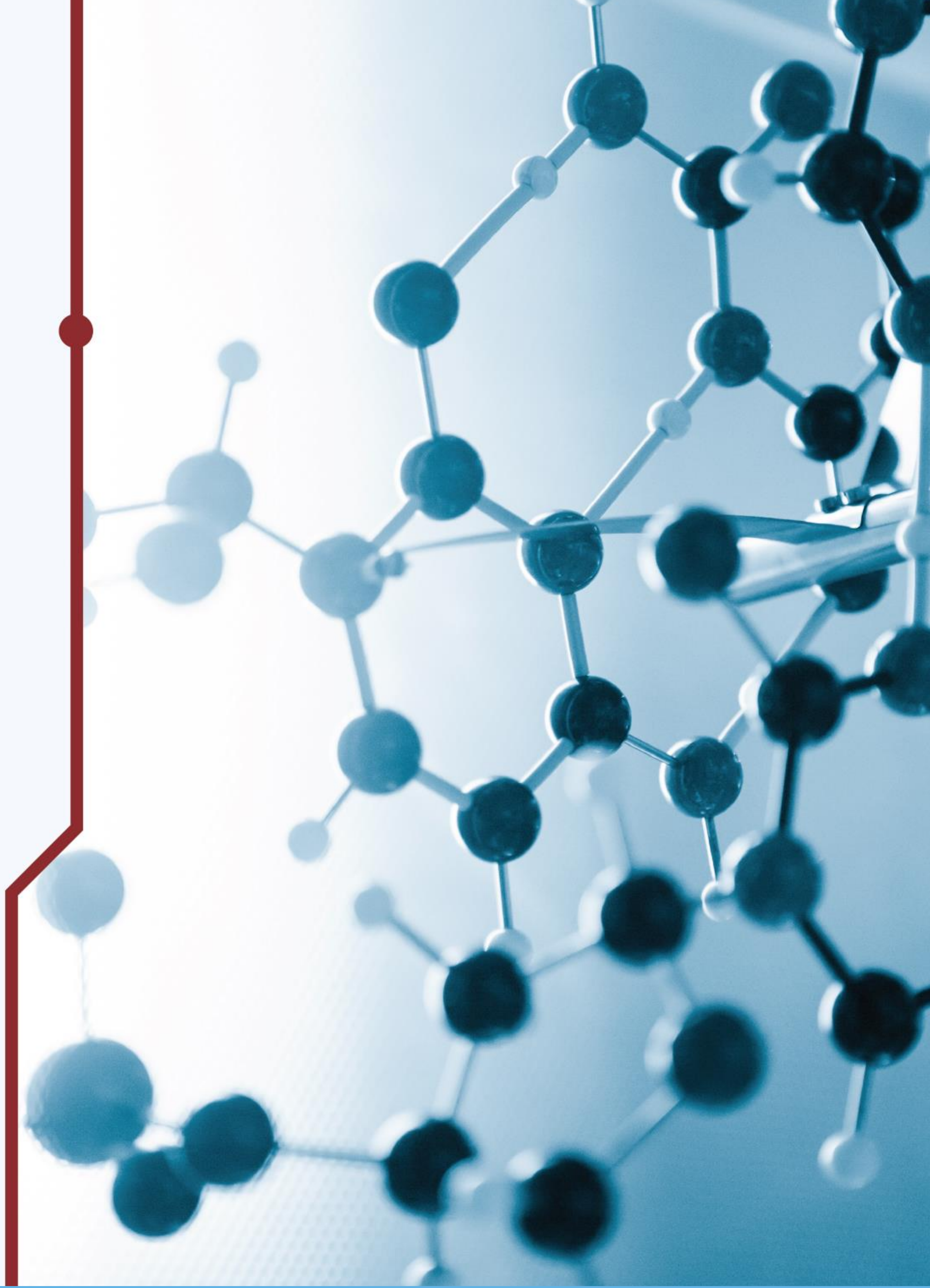
REsearch INfrastructures FOR Citizens in Europe



Gravitational Wave Noise Hunting

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On behalf of the REINFORCE-WP3

Reinforce Webinar

1 June 2020

The era of Gravitational Waves

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- **A new window on the Universe**
  - Study gravitational fields and mass distribution in cosmic sources
  - Probing black holes and other “dark” astrophysical sources
  - Test general relativity against other theories on gravitation
  - Investigate Big Bang cosmology (primordial gravitational waves)

- **Multimessenger Astrophysics**
  - Gravitational waves carry complementary information with respect of light (traditional astronomical “messenger”)

Gravitational Waves – a timeline

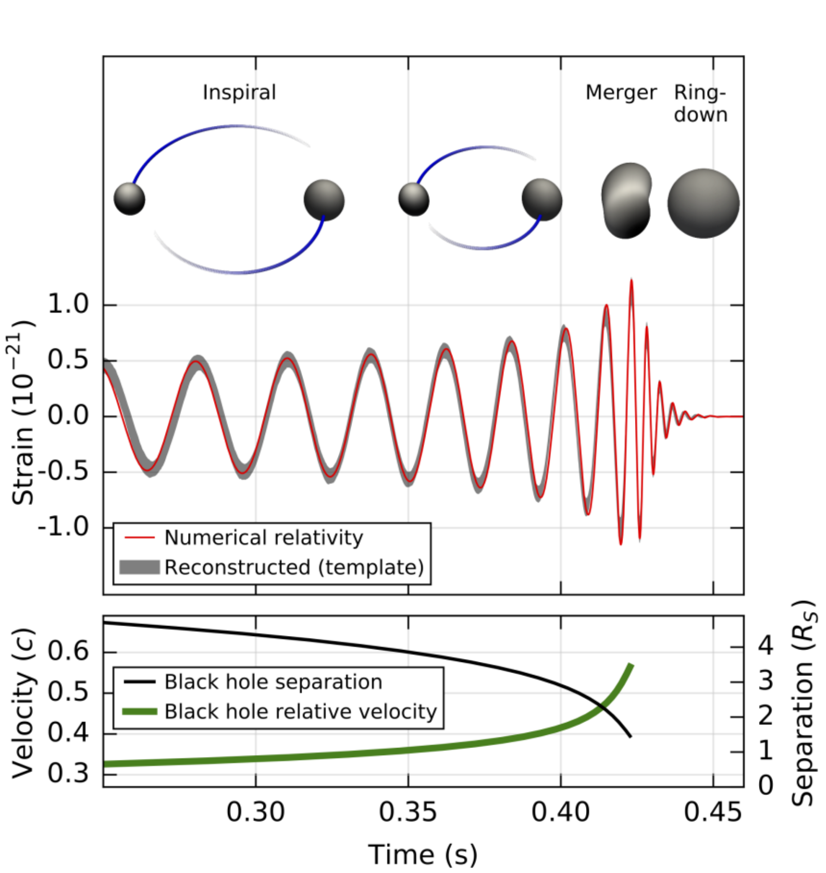
- **1915:** Einstein’s general relativity (new theory of gravity)
- **1916:** Einstein predict gravitational waves from general relativity
- **1968:** First attempts of detection by Joseph Weber (Maryland). Start of resonant bar projects
- **1972:** First tests on detectors based on interferometry (USA)
- **1981:** Start of studies in Italy on interferometry by Adalberto Giazotto
- **1984:** LIGO project funded in USA
- **1993:** Approval of Virgo project
- **1999:** Inauguration of LIGO detectors
- **2003:** Inauguration of Virgo detector
- **2007-2011:** Joint LIGO-Virgo observing runs
- **2011-2015:** Development of Advanced detectors (aiming at x10 sensitivity)
- **2015:** First detection of binary black hole GW (observing run O1)
- **2017:** Advanced Virgo joins LIGO in observing run O2. First detection of binary neutron star (17 Aug)
- **2019-2020:** Third observing run (end in 26 March 2020)

Sources of Gravitational Waves

- **What are gravitational waves?**
  - Ripples in spacetime traveling at the speed of light
  - Produced by acceleration or asymmetry of masses
  - Violent phenomena

- **Transient sources**
  - Inspiraling+merging of compact binary systems (black holes or neutron stars) – Detected!
  - Supernovae - expected
  - Others? - expected

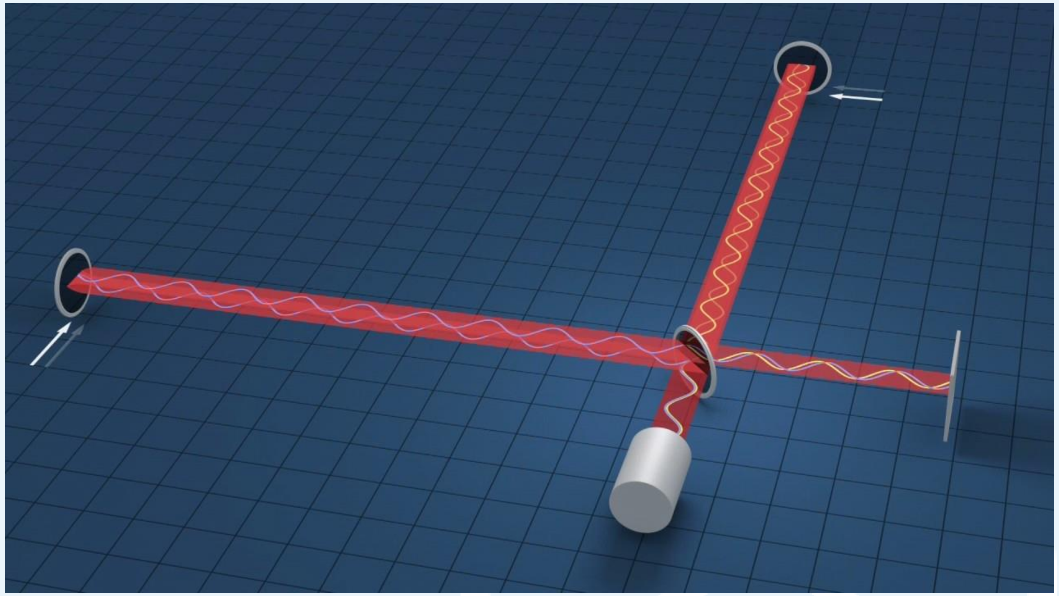
- **Continuous sources**
  - Periodic emission from rotating neutron stars (pulsars) – expected
  - Continuous stochastic background – expected
  - Others? - expected

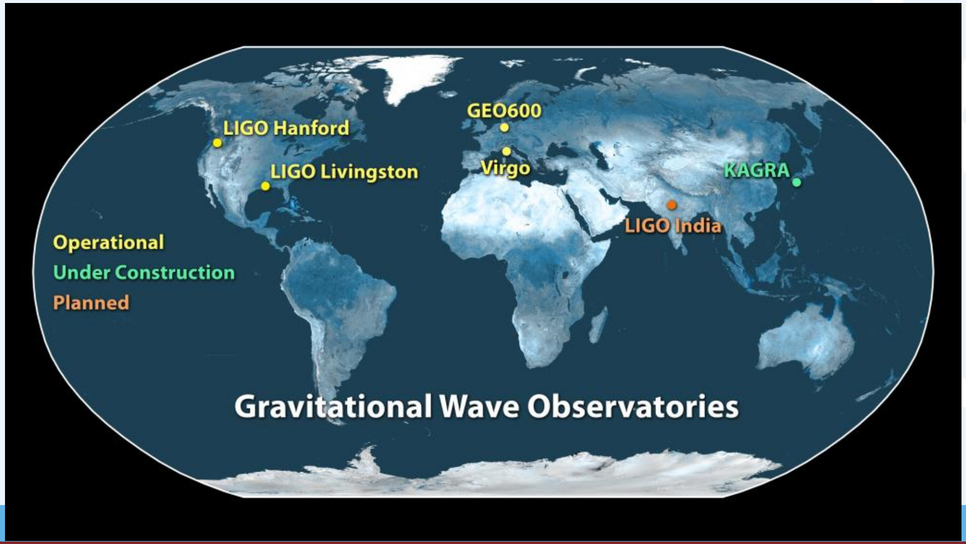


Abbott et al, 2016

Detecting Gravitational Waves

- **Extremely tiny signals**
  - A typical GW sources induce a deformation of 10^{-18} m over a length of ~1 km
  - High background noise
- **Laser interferometers**
  - Exploiting interference between orthogonal laser beams
  - Typical km-long scale
  - Frequency range 20-20000 Hz
  - Advanced methods to reduce noise
  - Detectors working as a network

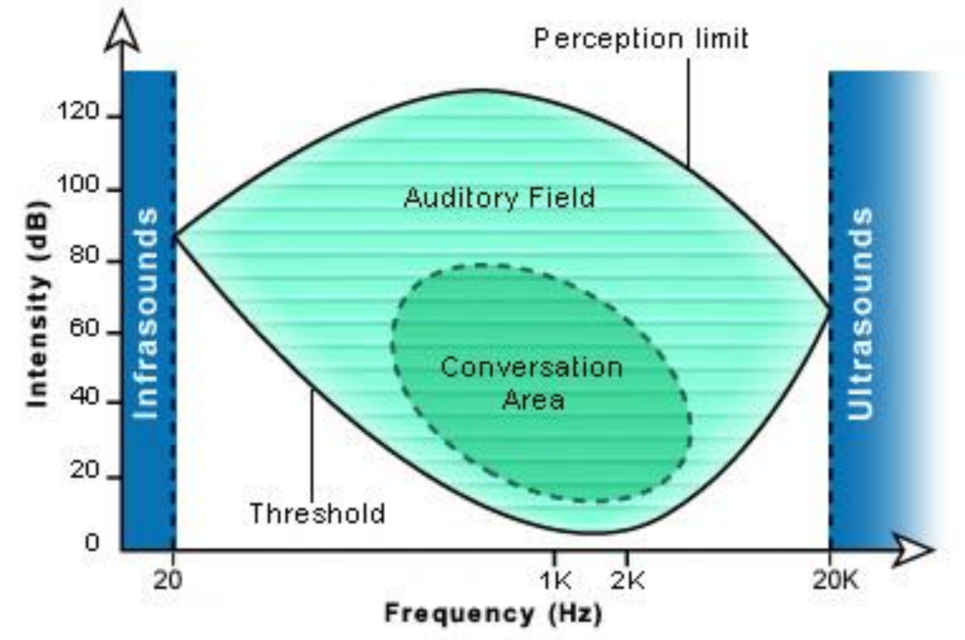


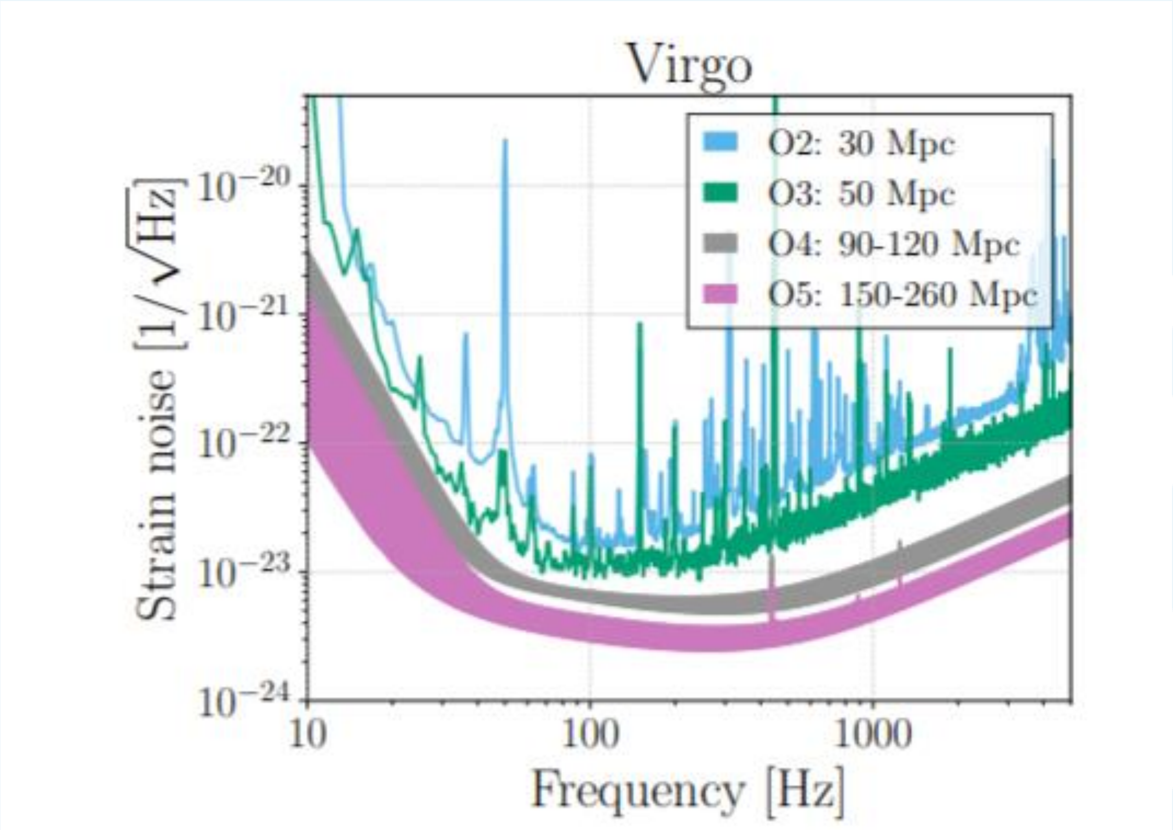


Credits: LIGO

Detecting Gravitational Waves

- **Sensitivity varies with frequency: main noise sources**
  - Low frequencies: Newtonian, seismic
  - Mid frequencies: thermal processes
  - High frequencies: quantum noise

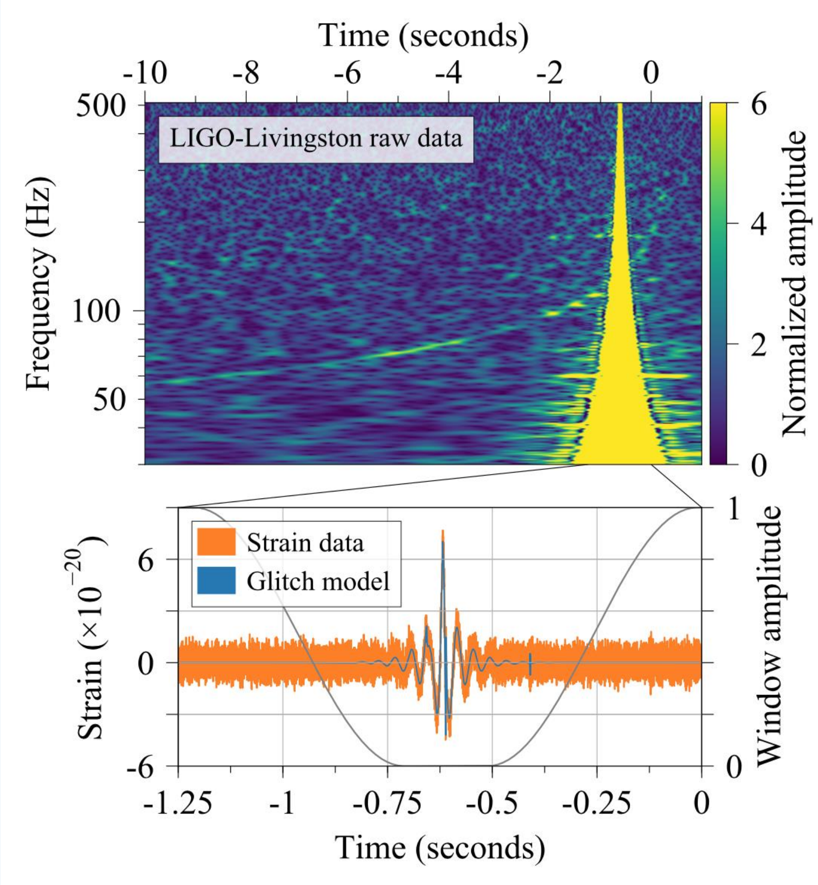




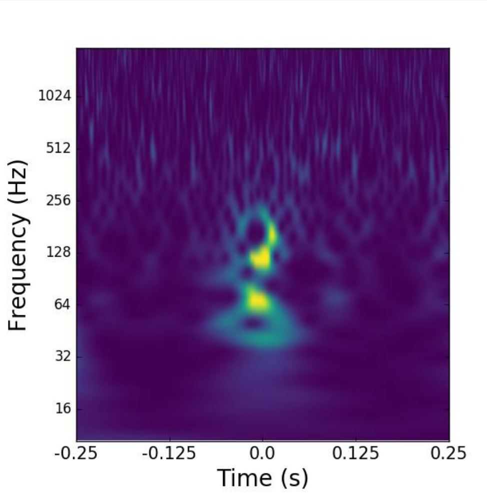
Abbott et al 2016

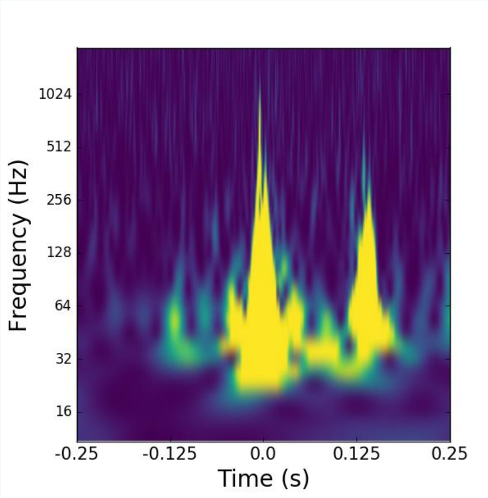
Noise glitches

- **Noise is not stationary in time**
  - Transient events can happen
  - Not related to astrophysical source, but local disturbances
  - They affect data quality and detection
- **Noise hunting & characterization is critical**
  - Detect and classify glitches to find origin and remove them
  - Glitches have complex morphology
  - Machine learning show promising approaches



Glitch in LIGO L1 detector during GW170817
Abbott et al 2017

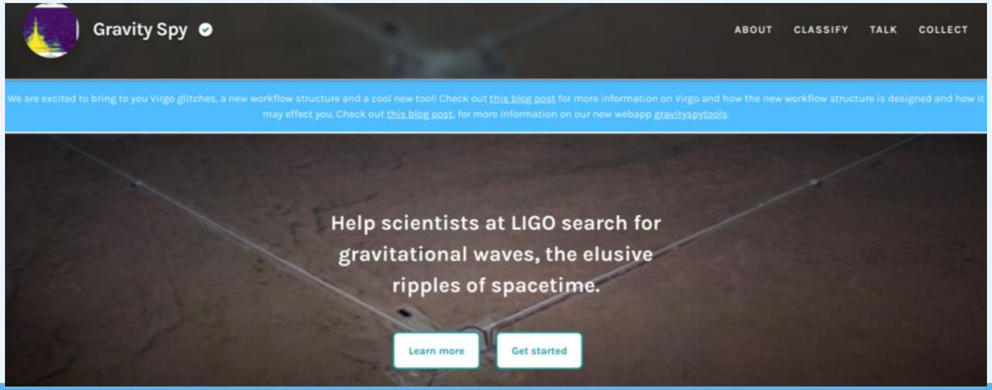


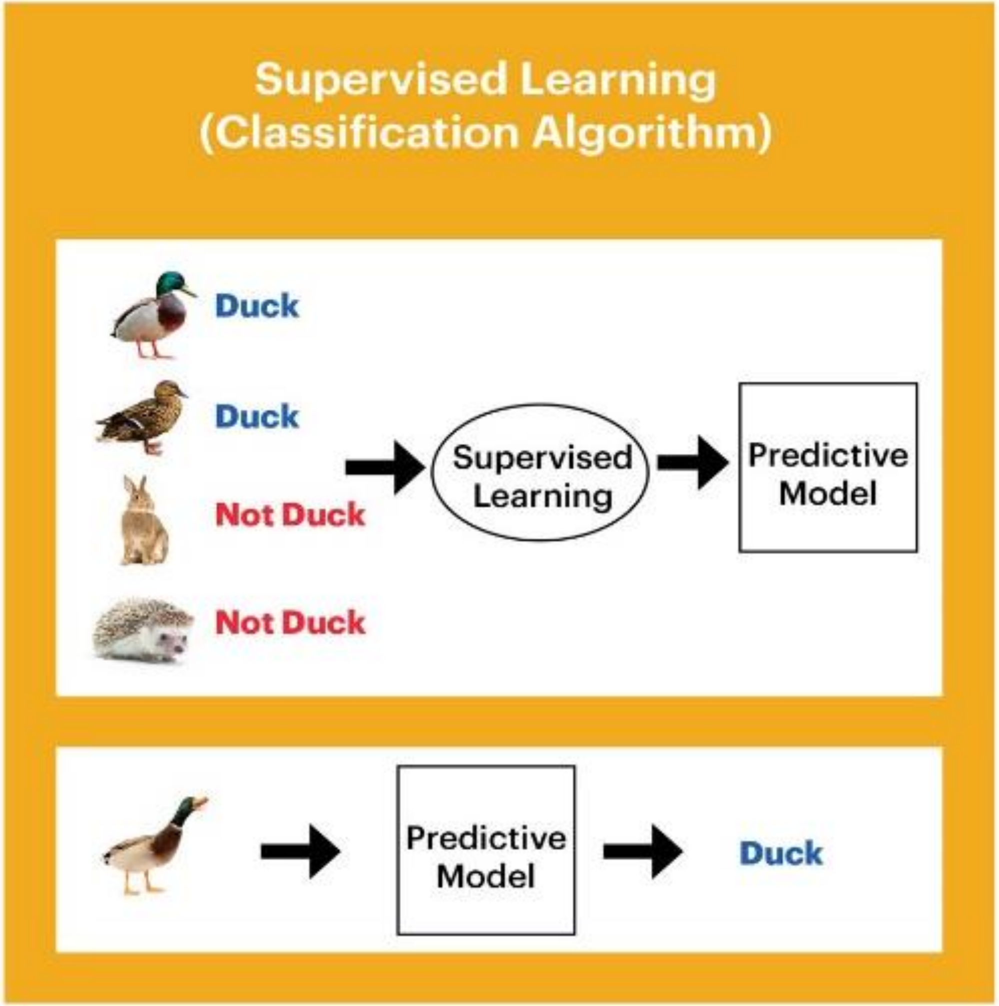


Glitches & citizen science

- **Machine learning approach**
  - Promising to classify complex time-frequency patterns of glitches
  - Large input to train machine learning models
  - Input from citizen science can be very important

- **Citizen scientists can help!**
  - Detect and classify glitches to find origin and remove
  - Glitches have complex morphology
  - Machine learning show promising approaches
  - Success story: Gravity Spy on ZooUniverse (2016)
  - see previous presentation





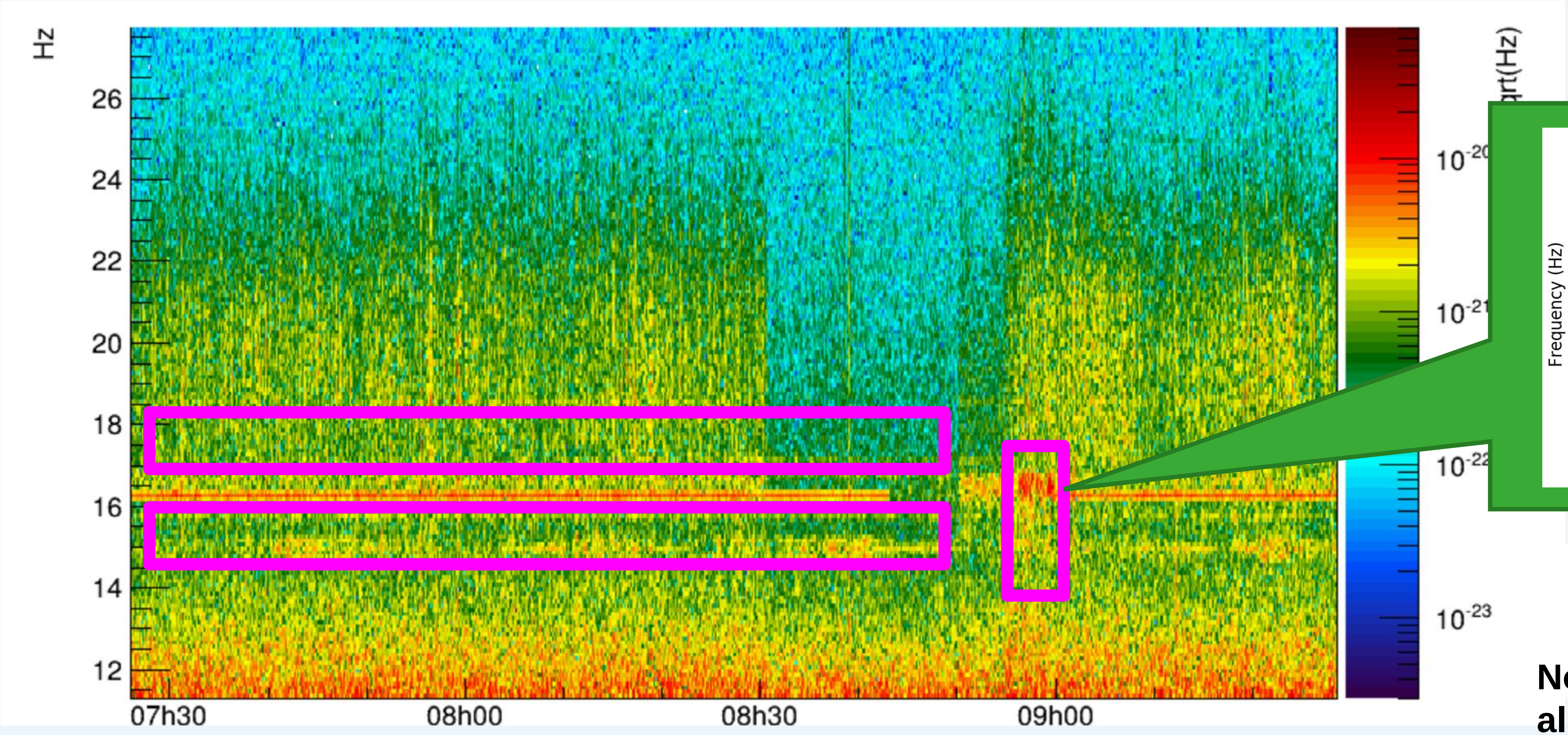
Credits: Western Digital

GW noise hunting in REINFORCE

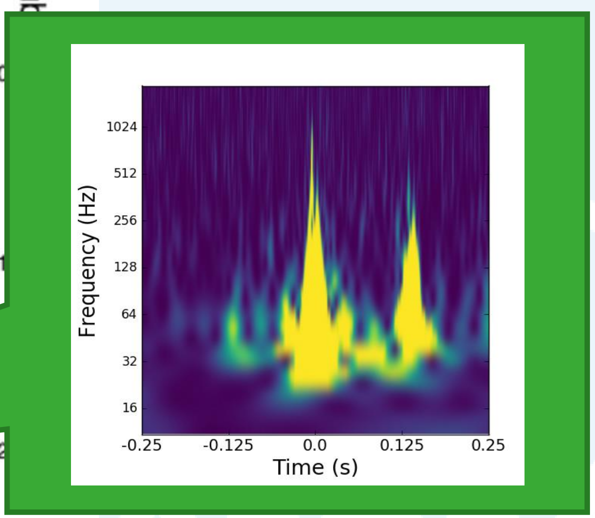
- **Noise hunting & citizen science**
  - Citizens can contribute to noise identification and classification
  - A specific project is under development as a part of REINFORCE
  - Demonstrator is the goal of work package 3
  - Unipi, EGO, EA, CONICET, OU, UOXF

- **Share findings and find new features**
  - Will use real data from GW detectors (LIGO, Virgo)
  - Not only known noise features. New one will be presented and citizens will suggest new, not yet labeled, noise types
  - Data will be presented via time-frequency representations

Investigating the noise



From Virgo logbook



Example of “zoom” on a glitch

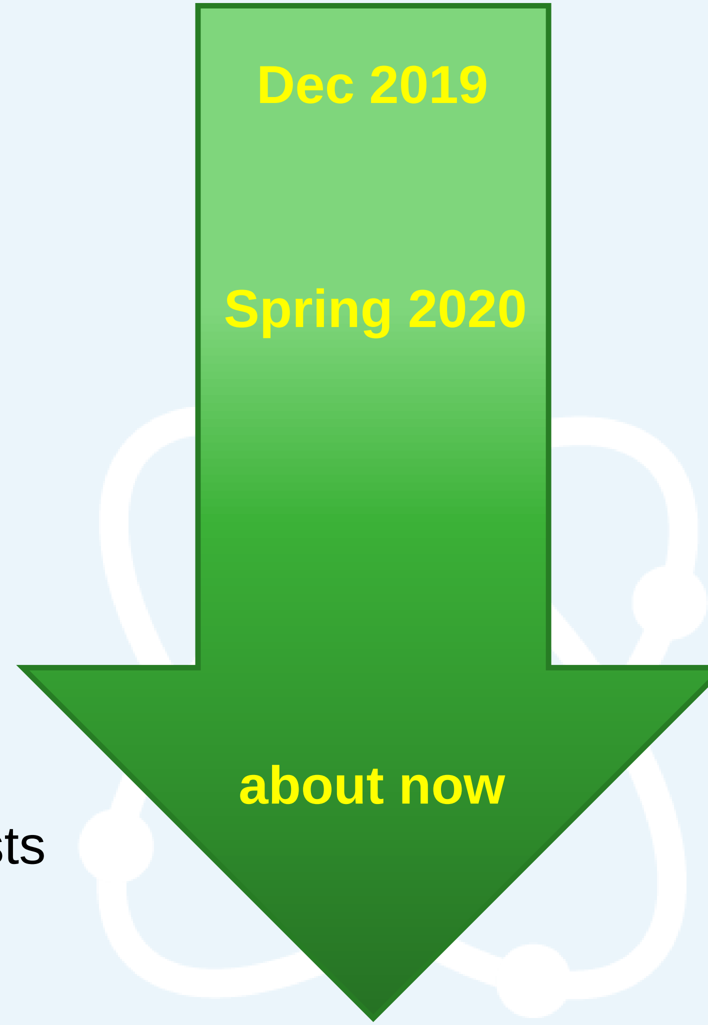
Not just visual: also audio representation will be used

The road ahead - I

- **Start !**

- **Establish data selection & Format**
  - Select data from LIGO and Virgo stream
  - Filter good time intervals of data
  - Develop flexible data format
  - Identify optimal visual and audio filtering

- **Dataset creation**
  - Select a statistically significative representation of the full data
  - Create visual and sound datasets that will be presented to citizen scientists



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The road ahead - II

- **Developing the ML algorithms**
  - Machine learning for classification
  - Identify optimal configuration for classification of noise starting from citizens' input

- **Developing and deploying the website**
  - Will be on Zoouniverse
  - Prepare documentation, guides
  - Open communication channels (e.g. blog, social)

- **Launch the website**

- **Comparative analysis**
  - Performance of human vs machine learning
  - Sound vs visual representation
  - Impact on science data

Summer 2020

Fall 2020

Summer 2021

Summer 2022

Conclusions

- Gravitational wave physics is a new, evolving field of science
- Big amount of data, contribution in analysis is welcome!
- Understanding background noise is a critical issue
- Less noise → More sensitivity → More events → More science!
- New GW demonstrator ready for launch in ~1 year from now